

# Nonnegative Matrix Factorization and Applications

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# Linear Inverse Problem

- Solve

$$Kx = y$$

in discrete setting

- $x \in \mathbb{R}^p$  = vector of coefficients describing the unknown object
- $y \in \mathbb{R}^n$  = vector of (noisy) data
- $K$  = linear operator ( $n \times p$  matrix) modelling the link between the two

# Regularization

Noisy data  $\rightarrow$  solve approximately by minimizing contrast (discrepancy) function, e.g.  $\|Kx - y\|_2^2$

Ill-conditioning  $\rightarrow$  regularize by adding constraints/penalties on the unknown vector  $x$  e.g.

- on its squared  $L^2$ -norm  $\|x\|_2^2 = \sum_i |x_i|^2$   
(classical quadratic regularization)
- on its  $L^1$ -norm  $\|x\|_1 = \sum_i |x_i|$   
(sparsity-enforcing or “lasso” regularization, favoring the recovery of sparse solutions, i.e. the presence of many zero components in  $x$ )
- on a linear combination of both  $\|x\|_1$  and  $\|x\|_2^2$  norms  
(“elastic-net” regularization)
- plus here **positivity constraints** (hold true in many applications)

# Positivity and multiplicative iterative algorithms

- Poisson noise  $\rightarrow$  minimize (log-likelihood) cost function subject to  $x \geq 0$  (assuming  $K \geq 0$  and  $y \geq 0$ )

$$F(x) = KL(y, Kx) \equiv \sum_{i=1}^n \left[ y_i \ln \left( \frac{y_i}{(Kx)_i} \right) - y_i + (Kx)_i \right]$$

(Kullback-Leibler – generalized – divergence)

- [Richardson \(1972\)](#) - [Lucy \(1974\)](#) (an astronomer's favorite) = EM(ML) in medical imaging

- Algorithm: 
$$x^{(k+1)} = \frac{x^{(k)}}{K^T \mathbf{1}} \circ K^T \frac{y}{Kx^{(k)}} \quad (k = 0, 1, \dots)$$

(using the Hadamard (entrywise) product  $\circ$  and division;  
 $\mathbf{1}$  is a vector of ones)

- Positivity automatically preserved if  $x^{(0)} > 0$
- Unregularized  $\rightarrow$  semi-convergence  $\rightarrow$  usually early stopping
- Can be easily derived through separable surrogates

# Surrogating

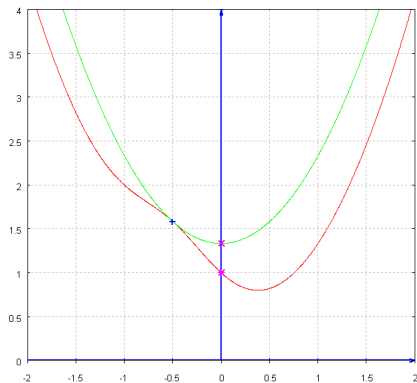


Figure : The function in red and his surrogate in green

# Surrogating

- Surrogate cost function  $G(x; a)$  for  $F(x)$ :

$$G(x; a) \geq F(x) \quad \text{and} \quad G(a; a) = F(a)$$

for all  $x, a$

- MM-algorithm (Majorization-Minimization):

$$x^{(k+1)} = \arg \min_x G(x; x^{(k)})$$

- Monotonic decrease of the cost function is then ensured:

$$F(x^{(k+1)}) \leq F(x^{(k)})$$

(Lange, Hunter and Yang 2000)

# Surrogate for Kullback-Leibler

Cost function ( $K \geq 0$  and  $y \geq 0$ )

$$F(x) = \sum_{i=1}^n \left[ y_i \ln \left( \frac{y_i}{(Kx)_i} \right) - y_i + (Kx)_i \right]$$

Surrogate cost function (for  $x \geq 0$ )

$$G(x; a) = \sum_{i=1}^n \left[ y_i \ln y_i - y_i + (Kx)_i + \right. \\ \left. - \frac{y_i}{(Ka)_i} \sum_{j=1}^p K_{i,j} a_j \ln \left( \frac{x_j}{a_j} (Ka)_i \right) \right]$$

NB. This surrogate is separable, i.e. it can be written as a sum of terms, where each term depends only on a single unknown component  $x_j$ .

# Positivity and multiplicative iterative algorithms

- Gaussian noise  $\rightarrow$  minimize (log-likelihood) cost function subject to  $x \geq 0$

$$F(x) = \frac{1}{2} \|Kx - y\|_2^2$$

assuming  $K \geq 0$  and  $y \geq 0$

- ISRA (Image Space Reconstruction Algorithm)  
(Daube-Witherspoon and Muehllehner 1986; De Pierro 1987)
- Iterative updates

$$x^{(k+1)} = x^{(k)} \circ \frac{K^T y}{K^T K x^{(k)}}$$

- Positivity automatically preserved if  $x^{(0)} > 0$
- Unregularized  $\rightarrow$  semi-convergence  $\rightarrow$  usually early stopping
- Easily derived through separable surrogates



# Surrogate for Least Squares

Cost function ( $K \geq 0$  and  $y \geq 0$ )

$$F(x) = \frac{1}{2} \|Kx - y\|_2^2$$

Surrogate cost function (for  $x \geq 0$ )

$$G(x; a) = \frac{1}{2} \sum_{i=1}^n \frac{1}{(Ka)_i} \sum_{j=1}^p K_{i,j} a_j \left[ y_i - \frac{x_j}{a_j} (Ka)_i \right]^2$$

NB. This surrogate is separable, i.e. it can be written as a sum of terms, where each term depends only on a single unknown component  $x_j$

# Blind Inverse Imaging

- In many instances, the operator is unknown (“blind”) or only partially known (“myopic” imaging/deconvolution)
- The resulting functional is convex w.r.t.  $x$  or  $K$  separately but is **not jointly convex**  $\rightarrow$  possibility of local minima
- Usual strategy: alternate minimization on  $x$  (with  $K$  fixed) and  $K$  (with  $x$  fixed)
- The problem can be easily generalized to include multiple inputs/unknowns ( $x$  becomes a  $p \times m$  matrix  $X$ ) and multiple outputs/measurements ( $y$  becomes a  $n \times m$  matrix  $Y$ ) e.g. for Hyperspectral Imaging

$$\longrightarrow \quad \text{solve} \quad KX = Y$$

## Special case: Blind Deconvolution

- When the imaging operator  $K$  is translation-invariant, the problem is also referred to as “Blind Deconvolution”
- Alternating minimization approaches using (regularized) least-squares (Ayers and Dainty 1988; You and Kaveh 1996; Chan and Wong 1998, 2000) or Richardson-Lucy (Fish, Brinicombe, Pike and Walker 1996)
- Bayesian approaches are also available
- An interesting non-iterative and nonlinear inversion method has been proposed by Justen and Ramlau (2006) with a uniqueness result. Unfortunately, their solution has been shown to be unrealistic from a physical point of view by Carasso (2009)

# Blind Inverse Imaging, Positivity and NMF

- Blind imaging is difficult → use as much a priori information and constraints as you can
- In particular, positivity constraints have proved very powerful when available, e.g. in incoherent imaging as for astronomical images
- The special case where all elements of  $K$ ,  $X$  (and  $Y$ ) are nonnegative ( $K \geq 0$ ,  $X \geq 0$ ) is also referred to as “Nonnegative Matrix Factorization” (NMF)
- There is a lot of recent activity on NMF, as an alternative to SVD/PCA for dimension reduction
- Alternating (ISRA or RL) multiplicative algorithms have been popularized by [Lee and Seung \(1999, 2000\)](#)  
See also [Donoho and Stodden \(2004\)](#)

# Our goal

- Develop a general and versatile framework for
- blind deconvolution/inverse imaging with positivity,
- equivalently for Nonnegative Matrix Factorization,
- with convergence results to control not only the decay of the cost function but also the convergence of the iterates,
- with algorithms simple to implement
- and reasonably fast...

We will consider

- Blind deconvolution with positivity  
(from single or multiple images)
- Hyperspectral Imaging
- Dynamic Positron Emission Tomography (PET)

NB. There are many other applications of NMF

## Regularized least-squares (Gaussian noise)

- Minimize the cost function, for  $K$ ,  $X$  nonnegative (assuming  $Y$  nonnegative too),

$$F(K, X) = \frac{1}{2} \|Y - KX\|_F^2 + \frac{\mu}{2} \|K\|_F^2 + \lambda \|X\|_1 + \frac{\nu}{2} \|X\|_F^2$$

where  $\|\cdot\|_F^2$  denotes the Frobenius norm  $\|K\|_F^2 = \sum_{i,j} K_{i,j}^2$

- The minimization can be done column by column on  $X$  and line by line on  $K$

# Regularized least-squares (Gaussian noise)

- Alternating multiplicative algorithm  
( $\mathbf{1}_{p \times m}$  is a  $p \times m$  matrix of ones)

$$K^{(k+1)} = K^{(k)} \circ \frac{Y(X^{(k)})^T}{K^{(k)}X^{(k)}(X^{(k)})^T + \mu K^{(k)}}$$

$$X^{(k+1)} = X^{(k)} \circ \frac{(K^{(k+1)})^T Y}{(K^{(k+1)})^T K^{(k+1)}X^{(k)} + \nu X^{(k)} + \lambda \mathbf{1}_{p \times m}}$$

- to be initialized with arbitrary but strictly positive  $K^{(0)}$  and  $X^{(0)}$
- Can be derived through surrogates  $\rightarrow$  provides a monotonic decrease of the cost function at each iteration
- Special cases:
  - a blind algorithm proposed by [Hoyer \(2002, 2004\)](#) for  $\mu = 0, \nu = 0$
  - ISRA for  $K$  fixed and  $\lambda = \mu = \nu = 0$



# Regularized least-squares (Gaussian noise)

- Assume  $\mu$  and either  $\nu$  or  $\lambda$  are strictly positive and  $Y$  has at least one strictly positive element in each row and each column
- Monotonicity is strict iff  $(K^{(k+1)}, X^{(k+1)}) \neq (K^{(k)}, X^{(k)})$
- The sequence  $F(K^{(k)}, X^{(k)})$  converges
- Asymptotic regularity holds:  $\forall i, j$   
$$\lim_{k \rightarrow +\infty} \left( K_{i,j}^{(k+1)} - K_{i,j}^{(k)} \right) = 0 ; \lim_{k \rightarrow +\infty} \left( X_{i,j}^{(k+1)} - X_{i,j}^{(k)} \right) = 0$$
- $\Rightarrow$  the set of accumulation points of the sequence of iterates  $(K^{(k)}, X^{(k)})$  is compact and connected
- If this set is finite, the iterates  $(K^{(k)}, X^{(k)})$  converge to a stationary point  $(K^*, X^*)$  (satisfying the first-order KKT conditions)


## Some recent related (methodological) work

(with convergence results, possibly positivity constraints)

- Algorithm based on the SGP algorithm by [Bonettini, Zanella, Zanni \(2009\)](#) and inexact block coordinate descent ([Bonettini 2011](#)): [Prato, La Camera, Bonettini, Bertero \(2013\)](#)  
For a space-variant PSF, see also [Ben Hadj, Blanc-Féraud and Aubert \(2012\)](#)
- Proximal Alternating Minimization and Projection Methods for Nonconvex Problems  
([Attouch, Bolte, Redont, Soubeyran 2010](#); [Bolte, Combettes and Pesquet 2010](#); [Bolte, Sabach and Teboulle 2013](#))
- Underapproximations for Sparse Nonnegative Matrix Factorization  
([Gillis and Glineur 2010](#))

# Application of NMF to Blind Deconvolution

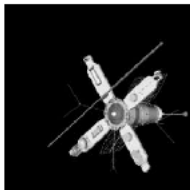
- $X$  :  $256 \times 256$  positive image
- $K$  : Convolution with the Airy function (circular low-pass filter)



The diagram illustrates the convolution equation  $Y = K * X$ . It consists of three square images arranged horizontally, separated by an equals sign and an asterisk. The first image, labeled  $Y$ , is a blurred version of the satellite image. The second image, labeled  $K$ , is a small white dot on a black background, representing the Airy function. The third image, labeled  $X$ , is the original sharp satellite image. The equation is  $Y = K * X$ .

# Application (Gaussian noise): no noise added

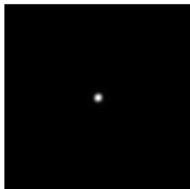
Original Image



Blurred Image



Reconstructed Psf



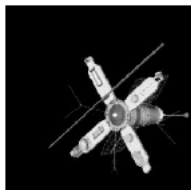
Reconstructed Image



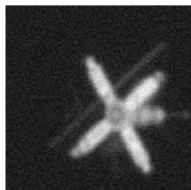
Figure :  $K^{(0)}$  Unif,  $X^{(0)} = \text{Blurred Image}$ ;  $\mu = 0$ ,  $\lambda = 0$ ,  $\nu = 0$ , 1000 it

# Application (Gaussian noise): 2.5% noise added

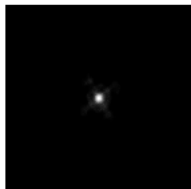
Original Image



Blurred and Noisy Image



Reconstructed PSF



Reconstructed Image

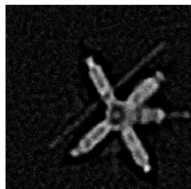


Figure :  $K^{(0)}$  Gaussian,  $X^{(0)}$  = Noisy Image;  $\mu = 2.25 \cdot 10^8$ ,  $\lambda = 0.03$ ,  $\nu = 0.008$ ;  
200 it

## Regularized Kullback-Leibler (Poisson noise)

- Minimize the cost function, for  $K$ ,  $X$  nonnegative (assuming  $Y$  nonnegative too),

$$F(K, X) = KL(Y, KX) + \frac{\mu}{2} \|K\|_F^2 + \lambda \|X\|_1 + \frac{\nu}{2} \|X\|_F^2$$

with

$$KL(Y, KX) = \sum_{i=1}^n \sum_{j=1}^m \left[ (Y)_{i,j} \ln \left( \frac{(Y)_{i,j}}{(KX)_{i,j}} \right) - (Y)_{i,j} + (KX)_{i,j} \right]$$

# Regularized Kullback-Leibler (Poisson noise)

- Alternating multiplicative algorithm

$$K^{(k+1)} = \frac{2A^{(k)}}{B^{(k)} + \sqrt{B^{(k)} \circ B^{(k)} + 4\mu A^{(k)}}}$$

where

$$A^{(k)} = K^{(k)} \circ \frac{Y}{K^{(k)} X^{(k)}} (X^{(k)})^T$$

$$B^{(k)} = \mathbf{1}_{n \times m} (X^{(k)})^T$$

( $\mathbf{1}_{n \times m}$  is a  $n \times m$  matrix of ones)

## Regularized Kullback-Leibler (Poisson noise)

$$X^{(k+1)} = \frac{2C^{(k+1)}}{D^{(k+1)} + \sqrt{D^{(k+1)} \circ D^{(k+1)} + 4vC^{(k+1)}}}$$

where

$$C^{(k+1)} = X^{(k)} \circ (K^{(k+1)})^T \frac{Y}{K^{(k+1)} X^{(k)}}$$

$$D^{(k+1)} = \lambda \mathbf{1}_{p \times m} + (K^{(k+1)})^T \mathbf{1}_{n \times m}$$

to be initialized with arbitrary but strictly positive  $K^{(0)}$  and  $X^{(0)}$



## Regularized Kullback-Leibler (Poisson noise)

- Can be derived through surrogates  $\rightarrow$  provides a monotonic decrease of the cost function at each iteration
- Special case for  $\lambda = \mu = \nu = 0$ : the blind algorithm proposed by [Lee and Seung \(1999\)](#) which reduces to the EM/Richardson-Lucy algorithm for  $K$  fixed
- Properties as above for the least-squares case

# Normalization constraint

- At each iteration, one can enforce a normalization constraint on the PSF (line of  $K$ ), imposing that its values sum to one
- To do this a Lagrange multiplier is introduced and its value is determined by means of a few Newton-Raphson iterations
- The convergence results can be adapted to cope with this case

# Application : 1% (equiv. rmse) Poisson Noise; PSF normalized

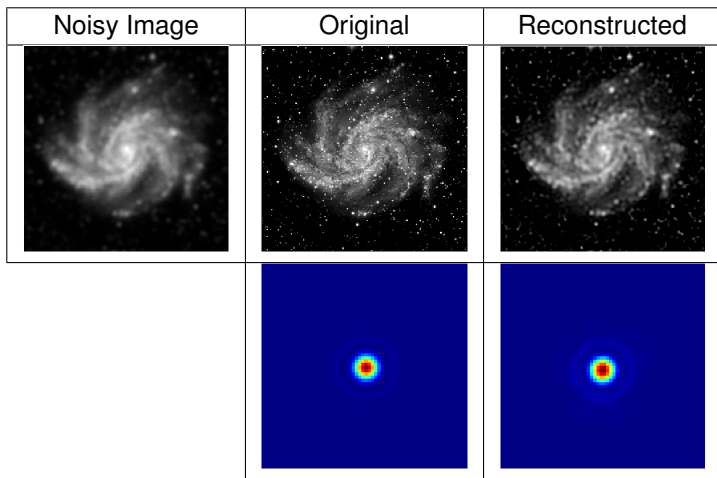


Figure :  $K^{(0)} = \text{Unif}$ ,  $X^{(0)} = \text{Noisy Image}$ ,  $\mu = 10^9$ ,  $\lambda = 10^{-7}$ ,  $\nu = 6 \cdot 10^{-8}$ ,  
2000 it in 12m37s

# Extension to TV regularization

- Total Variation: use discrete differentiable approximation

$$\|X\|_{TV} = \sum_{i,j} \sqrt{\varepsilon^2 + (X_{i+1,j} - X_{i,j})^2 + (X_{i,j+1} - X_{i,j})^2}$$

for 2D images

- Use penalty  $\lambda\|X\|_{TV}$  instead of  $\lambda\|X\|_1$
- Use separable surrogate proposed by [Defrise, Vanhove and Liu \(2011\)](#) to derive explicit update rules both for Gaussian and Poisson noise

# Application KL-TV: 1% (equiv. rmse) Poisson Noise; PSF normalized

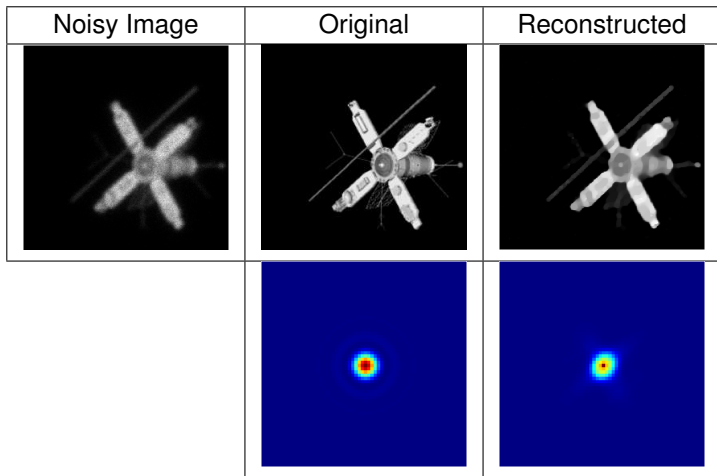


Figure :  $K^{(0)} = \text{Unif}$ ,  $X^{(0)} = \text{Noisy Image}$ ,  $\mu = 1.5 \cdot 10^6$ ,  $\lambda = 0.0485$ ,  
 $\varepsilon = 6 \cdot 10^{-7}$ , 200 it in 1m46s

# Application KL-TV: 2.5% (equiv. rmse) Poisson Noise; normalized PSF

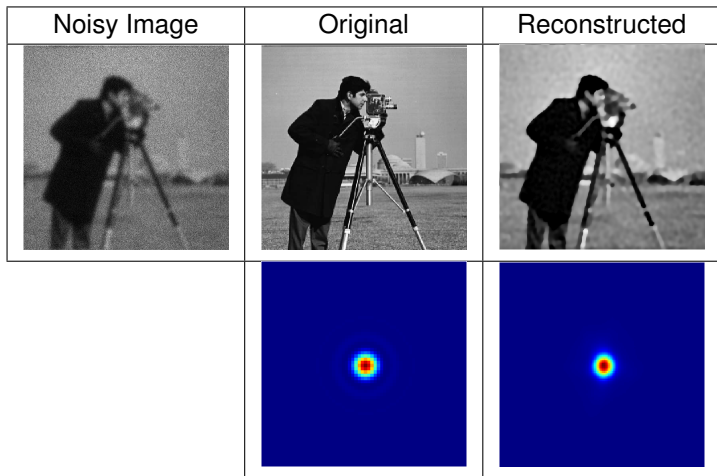


Figure :  $K^{(0)} = \text{Unif}$ ,  $X^{(0)} = \text{Noisy Image}$ ,  $\mu = 10^7$ ,  $\lambda = 0.03$ ,  $\varepsilon = \sqrt{10}$ , 2000 it in 54m30s

# Application to AO astronomical images

- Test images used by [Prato, La Camera, Bonettini and Bertero \(2013\)](#) for post-adaptive-optics astronomical imaging (PSF with Strehl Ratio 0.40), with a restoration method based on the SGP algorithm by [Bonettini, Zanella, Zanni \(2009\)](#) and inexact block coordinate descent ([Bonettini 2011](#))
- Necessity to incorporate a background term in the cost function ( $KX \rightarrow KX + B$ ) and in the KL restoration algorithm
- Quality of restorations comparable to those of [Prato, La Camera, Bonettini and Bertero \(2013\)](#)

# Planetary Nebula NGC 7027

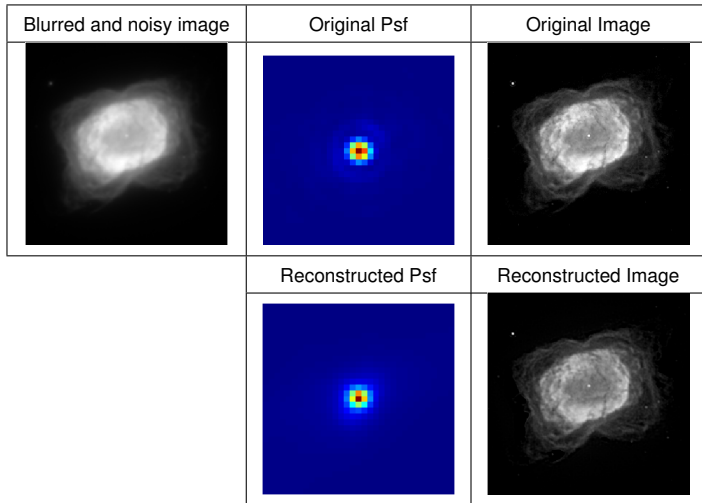


Figure :

Reconstruction with  $\mu = 10^6$ ,  $\lambda = 6 \cdot 10^{-8}$ ,  $\nu = 10^{-9}$ , uniform  $K^{(0)}$ ,  $X^{(0)} = Y$ ,  
5500 it in 17min40s,  $RMSE_K = 22.61\%$ ,  $RMSE_X = 6.66\%$



# Galaxy NGC 6946

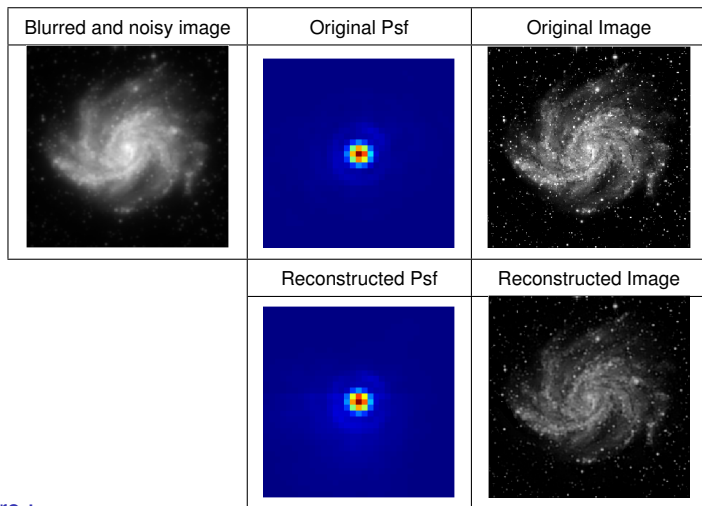


Figure :

Reconstruction with  $\mu = 2.75 \cdot 10^8$ ,  $\lambda = 5 \cdot 10^{-8}$ ,  $\nu = 1.668 \cdot 10^{-8}$ , uniform  $K^{(0)}$ ,  $X^{(0)} = Y$ ,  
5000 it in 15min45s,  $RMSE_K = 14.24\%$ ,  $RMSE_X = 22.83\%$

# Crab Nebula NGC 1952

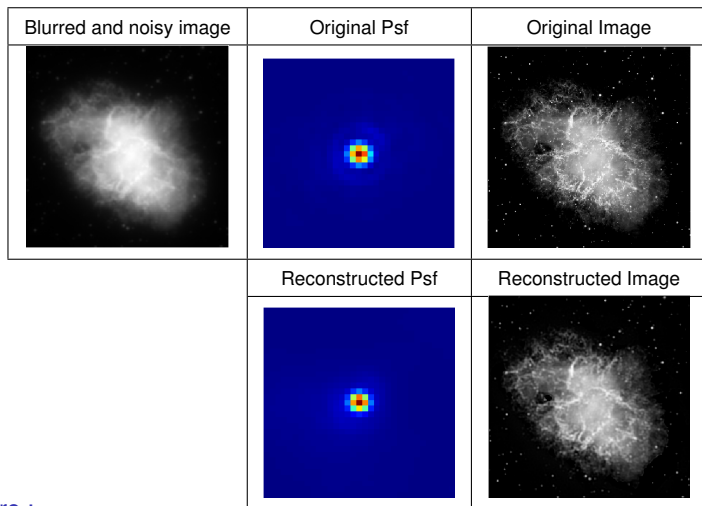



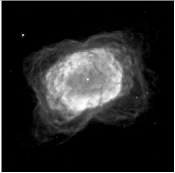
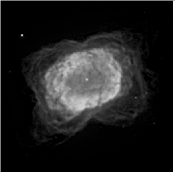
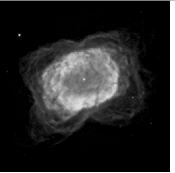

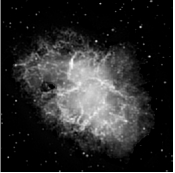



Figure :

Reconstruction with  $\mu = 7.4 \cdot 10^7$ ,  $\lambda = 2.5 \cdot 10^{-8}$ ,  $\nu = 1.5 \cdot 10^{-8}$ , uniform  $K^{(0)}$ ,  $X^{(0)} = Y$ ,  
4000 it in 11min48s,  $RMSE_K = 15.72\%$ ,  $RMSE_X = 15.96\%$

# Reconstruction from multiple images (same PSF)

	Original	Single Image	Multiple (3) Images
Galaxy			
Planetary Nebula			
Crab Nebula			

# Reconstruction from multiple images (same PSF)

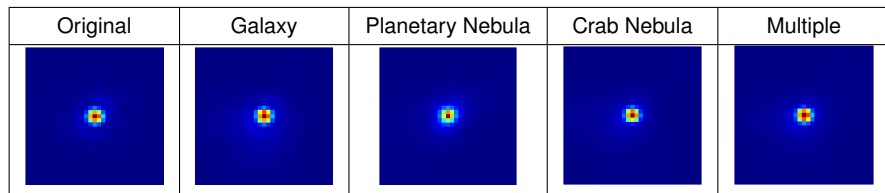


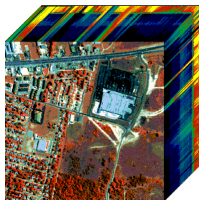
Figure : Reconstructed PSF

Object	$\mu$	$\lambda$	$\nu$	Nb It.	Time	$RMSE_K$	$RMSE_X$	$RMSE_M$
Plan.	$10^6$	$10^{-8}$	$10^{-9}$	5500	17m40s	0.23	0.06	0.05
Gal.	$2.75 \cdot 10^8$	$5 \cdot 10^{-8}$	$1.67 \cdot 10^{-8}$	5000	15m45s	0.14	0.23	0.19
Crab	$7.4 \cdot 10^7$	$2.5 \cdot 10^{-8}$	$1.5 \cdot 10^{-8}$	4000	11m48s	0.15	0.16	0.14
All 3	$2.4 \cdot 10^8$	$6.75 \cdot 10^{-8}$	$4.36 \cdot 10^{-9}$	4000	21m27s	0.12		

Table : Values used for the parameters and RMSE

# Application of NMF to Hyperspectral Imaging

Example: Urban HYDICE HyperCube:  $307 \times 307 \times 162$   
containing the images of an urban zone recorded for 162 different  
wavelength/frequencies



- Factorize the  $Y : 307^2 \times 162$  data matrix as  $Y = KX$  where  $K$  is a  $307^2 \times p$  (relative) abundances matrix of some basis elements to be determined and  $X$  is a  $p \times 162$  matrix containing the spectra of those basis elements
- Penalized Kullback-Leibler divergence used as cost function
- The sum of the relative abundances is normalized to one

# Hyperspectral Imaging

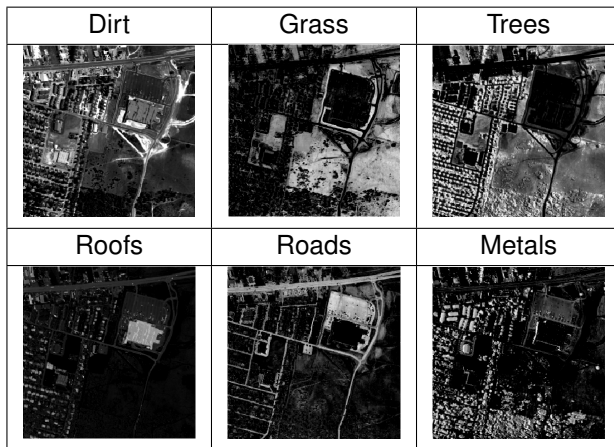


Figure :

Abundances with

$\rho = 6$ ,  $\mu = 10^{-10}$ ,  $\lambda = 0$ ,  $\nu = 1.1$ ,  
random  $K^{(0)}$  and  $X^{(0)}$ , 1000 it in 1h19min12s

# Hyperspectral Imaging

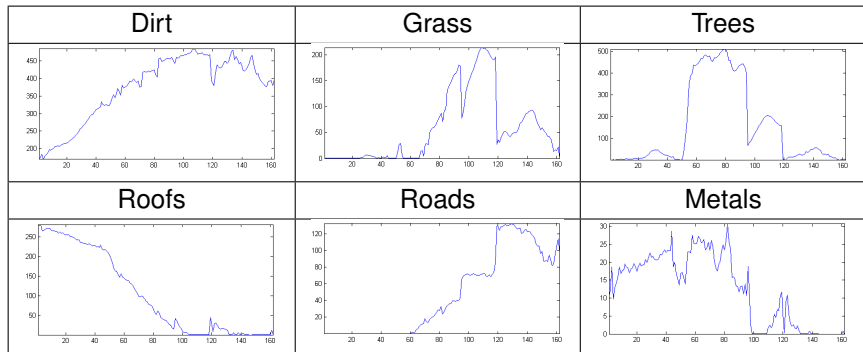
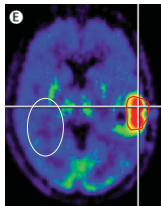
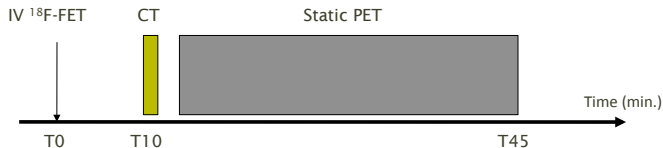


Figure : Spectra

# Application of NMF to Dynamic PET

(work in progress with Michel Defrise)

$^{18}\text{F}$ -Fluoro-ethyl-tyrosine ( $^{18}\text{F}$ -FET) =  
artificial amino-acid for PET, crossing the Blood Brain Barrier

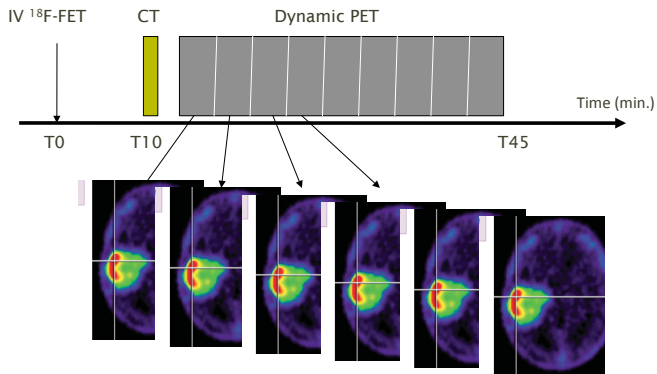


Pixels with FET-uptake  $\geq 1.6$  of normal brain are considered "tumor associated"

Sum of all tumor associated pixels:  
Metabolic Active Volume (MAV in ml)



# Dynamic PET



(slides : courtesy by Dr Hendrik Everaert)

NB. Large literature on NMF and in particular for dynamic PET, see e.g. Tichy, Smidl, 2014 International Conference on BioMedical Engineering and Informatics, Lee et al (Seoul) for myocardial  $\text{H}_2\text{O}$  PET, ...

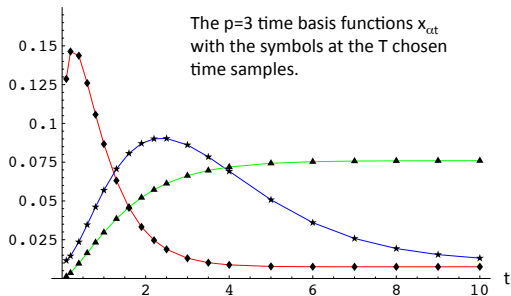
# Dynamic PET

PET scan data inversion is done time by time and then the 4D dataset (nb voxels  $\times$  nb of time frames) is factored through NMF, with cost function:

- Least-squares discrepancy (Gaussian noise)
- Smoothness of the temporal activity curves (TAC)/factors: quadratic penalty on a high-pass filtered version/differences
- Normalization of the temporal activity curves (TAC)/factors: sum to one over time
- Sparsity-enforcing penalty ( $L^1$ -norm) on the coefficients (few factors active in each voxel)
- Spatial correlations: quadratic penalty on differences of coefficients corresponding to neighbouring voxels

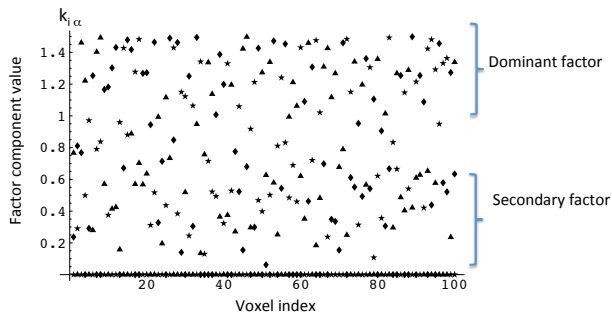
# Dynamic PET

A toy problem with 100 “voxels”, 20 time frames,  $p=3$  factors  
No spatial connection between voxels



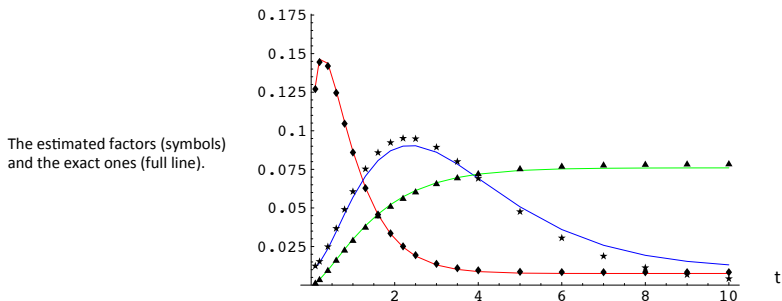
# Dynamic PET

The “true” coefficients  $k_{i\alpha}$  are selected as follows: each “voxel” has non-zero Coefficients for two factors randomly selected from the three factors. Given two random numbers  $r_1, r_2$  uniformly distributed on  $(0,1)$  the dominant coefficient is  $(0.5+\max(r_1,r_2))$  and the secondary factor modeling a weaker contamination is  $\min(r_1,r_2)$ . No spatial correlation is modeled.



# Dynamic PET

Toy problem: noise free simulation.  $\mu = \nu = 0$ ,  $N_{iter} = 10000$ .  $x_{ca}^{(0)} = 1$ ,  $k_{ii}^{(0)} = 0.1 + rand(0,1)$



The dominant factor  $\alpha_i^* = \operatorname{argmax}_{\alpha} k_{i\alpha}$  is identified correctly for all voxels.

The mean and standard deviation of the  $V=100$  errors on the coefficients are

Red factor: 0.01 (0.0069)

Blue factor: -0.05 (0.074)

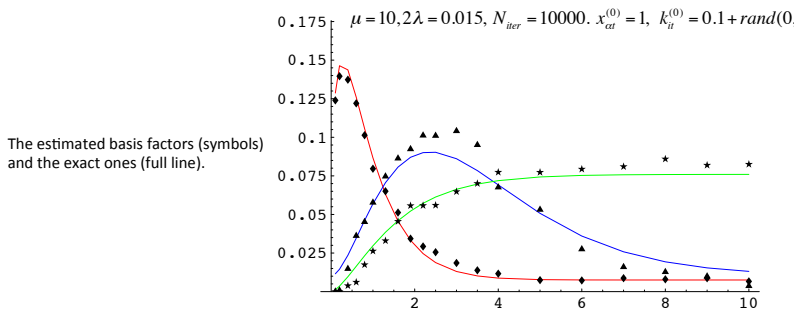
Green factor: 0.041 (0.074)

Relative RMSE of the reconstructed TAC (all time bins and all voxels):  $2.2 \cdot 10^{-6}$

Maximum TAC error/Maximum TAC: 0.00002.

# Dynamic PET

Toy problem: noisy simulation.



Despite noise and the important factor mixing, the dominant factor is identified correctly for 99 of the 100 voxels (for this noise realization).

The mean and standard deviation of the  $V=100$  errors on the coefficients are

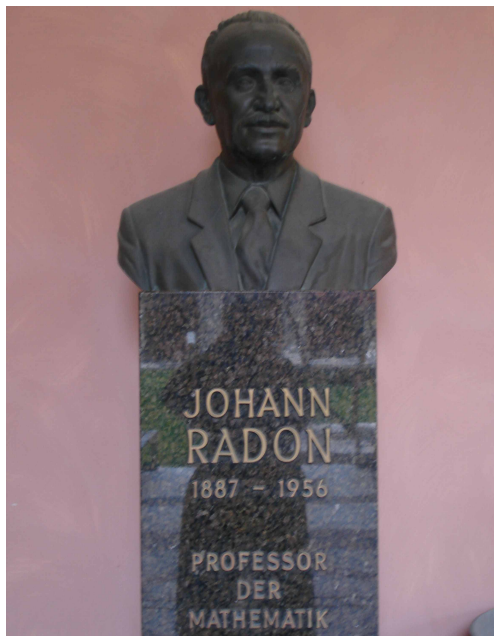
Red factor: -0.007 (0.10)

Blue factor: -0.14 (0.21)

Green factor: -0.15 (0.17)



# Happy Anniversary to the Radon Transform



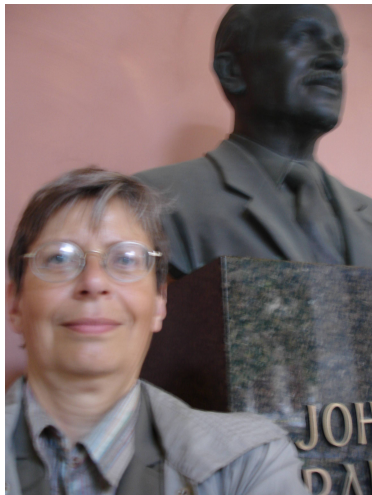
JOHANN  
RADON

1887 - 1956

PROFESSOR  
DER  
MATHEMATIK



A selfie from the pre-selfie era...



... and an invitation to blind deconvolution!